

Agenda: Kinetics and Transport in Multiparticle Systems

Dynamics of interacting multi-particle systems

- Interaction energies
Dissipation via multiple scattering
- Probabilistic evolution
Random walk and binomial distribution
Diffusion processes
Maxwell-Boltzmann energy distributions
Fluctuating (Langevin) dissipative forces
- Kinetics of dilute gases
Work and heat transfer
 - Flow of heat and radiation
 - Laws of thermodynamics, thermodynamic ensembles, entropy
- Fundamental ideal gas laws, Equation of state (EoS)

Reading Assignments

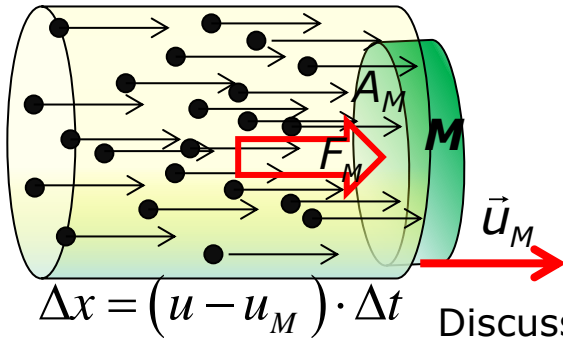
Weeks 5 & 6

LN IV.1-4:

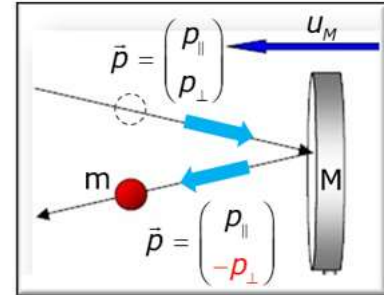
Kondepudi Ch. 1-3.
Additional Material

McQuarrie & Simon
Ch. 2, 5
Math Chapter B, C

Macroscopic Effects of Dissipative Energy Transfer



$$F_M = -2\rho_m \cdot A_M \cdot (u - u_M)^2$$



$$\Delta V = A_M \cdot \Delta x$$

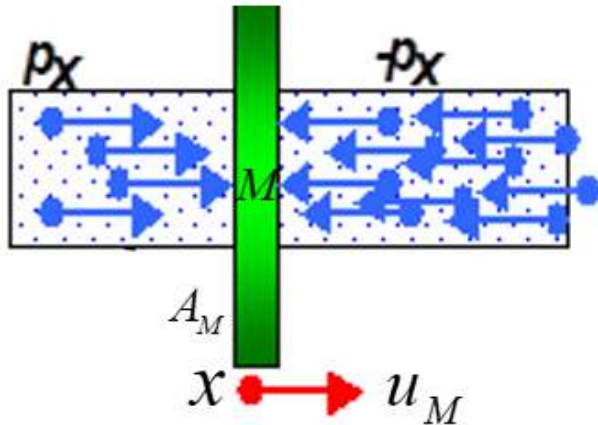
Discussion Points

Collision layer, Effects of multiple collisions with Mm
Trends to push M \rightarrow pV-work= conversion of random energy to directed flow/motion

Momentum mismatch with F/B ensembles, pos/neg p boost

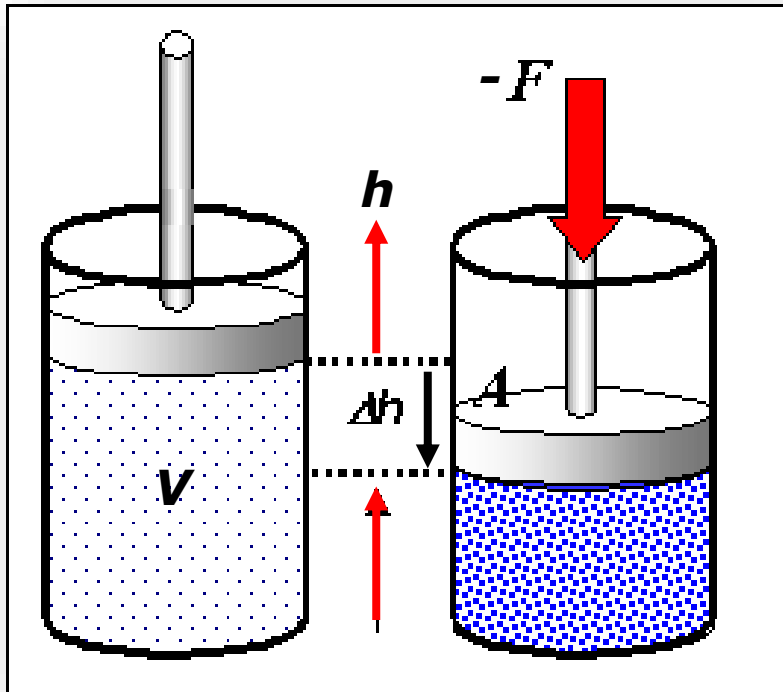
Exponents $\wedge 1$ vs. $\wedge 2$

Relation force-pressure @ $u_M=0$, ext vs. internal Expansion-cooling



$$\langle \underline{F_M} \rangle = -[4A_M m \int d\vec{p}^3 f(\vec{p}) \cdot |v_x|] \cdot \underline{u_M} = \underline{-\gamma \cdot u_M} \text{ with } \gamma > 0$$

Physical Changes: Compression of (Ideal) Gases



Compression of a gas **"System"** volume V with a constant force F (e.g., weight) on a constant area A :

→ Pressure $p = \text{Force } F / \text{Area } A$,
at $p = p_{ext} = \text{const. (external, not internal)}$

$$p = \frac{F}{A} = p_{ext} \rightarrow \Delta V = A \cdot \Delta h < 0$$

Compression work *done on* system

$$w = -F \cdot \Delta h = -(p \cdot A) \cdot \Delta h = -p \cdot \Delta V > 0$$

$$\text{Internal energy gain of gas: } \Delta U = -p \cdot \Delta V > 0 \rightarrow \Delta T > 0$$

Sign Convention work: Compressional work on a gas volume (=system) increases the internal energy E of the gas .

→ Work done on gas is counted as positive $w > 0()$.

Planck's Thermal (Blackbody) Radiation

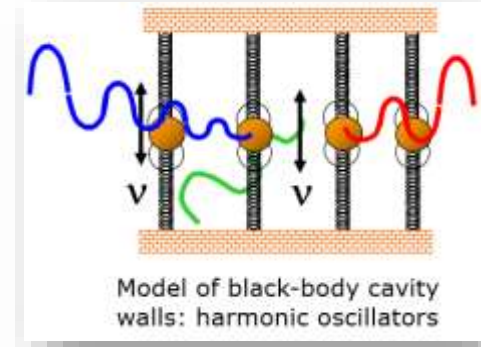


Experimental electromagnetic energy spectrum emitted by a mass ("black" body) @ T . Radiation (photon) energy

→ **Electromagnetic Energy Quantum**

$$E_\nu = h \cdot \nu = h \cdot c / \lambda$$

light vel. $c = \lambda \cdot \nu$



Planck's constant, from fit

$$h = 6.62618 \cdot 10^{-34} \text{ J} \cdot \text{s}$$

Power (energy/time) density@universal shape

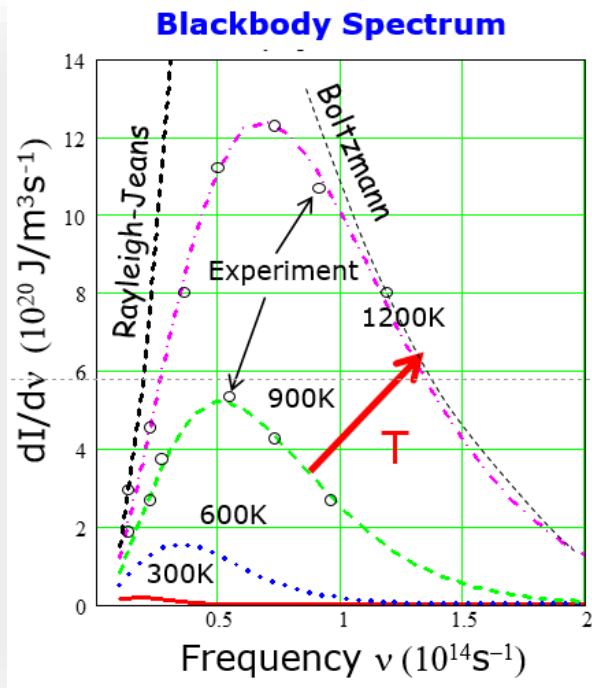
$$\frac{dI}{d\nu dV}(\nu, T) = \langle E_\nu \rangle \rho_\nu = \frac{8\pi}{c^3} \frac{h\nu^3}{e^{h\nu/k_B T} - 1}$$

Stefan-Boltzmann Law:

Total radiated thermal power

$$F = \int_0^\infty \frac{dI}{d\lambda \cdot dV} d\lambda = \sigma_{SB} \cdot T^4 \left(\text{W/m}^2 \right)$$

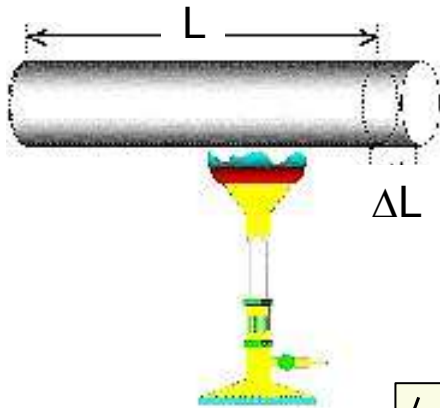
SB Constant $\sigma_{SB} = 5.67 \cdot 10^{-8} \frac{\text{W}}{\text{m}^2 \text{K}^4}$
 Def. "Watt" $W = J/s$



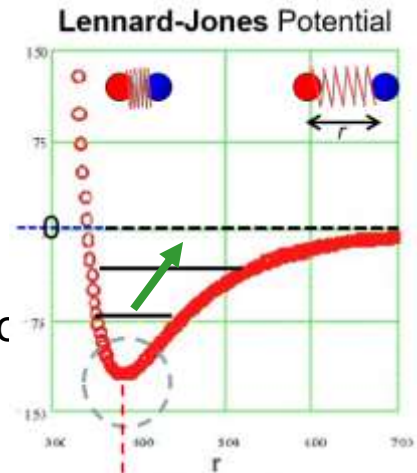
Conversion of Heat to Macroscopic Potential Energy

Heating a metal bar transfers energy to solid crystal lattice → bar expands in 3D. Reason: shape of the interaction potential between atoms/ions (Lennard-Jones-type or harmonic oscillator potential).

Macroscopic application: Avoid expansion effects by gaps between segments of bridges, highway concrete surface plates exposed to solar radiation



Effect of particle interaction potential $V(\text{distance}=r)$. Increased internal lattice energy weakens attractive forces → Increased atomic distance = macroscopic expansion.

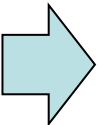


$$\langle L(T) \rangle = \langle L(T_0) \rangle [1 + \alpha (T - T_0)] > \langle L(T_0) \rangle$$

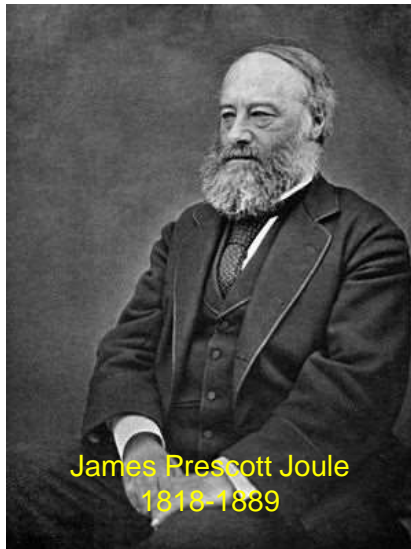
Linear expansion coefficient $\alpha > 0$

Can be anisotropic in 3D (lattice structure)

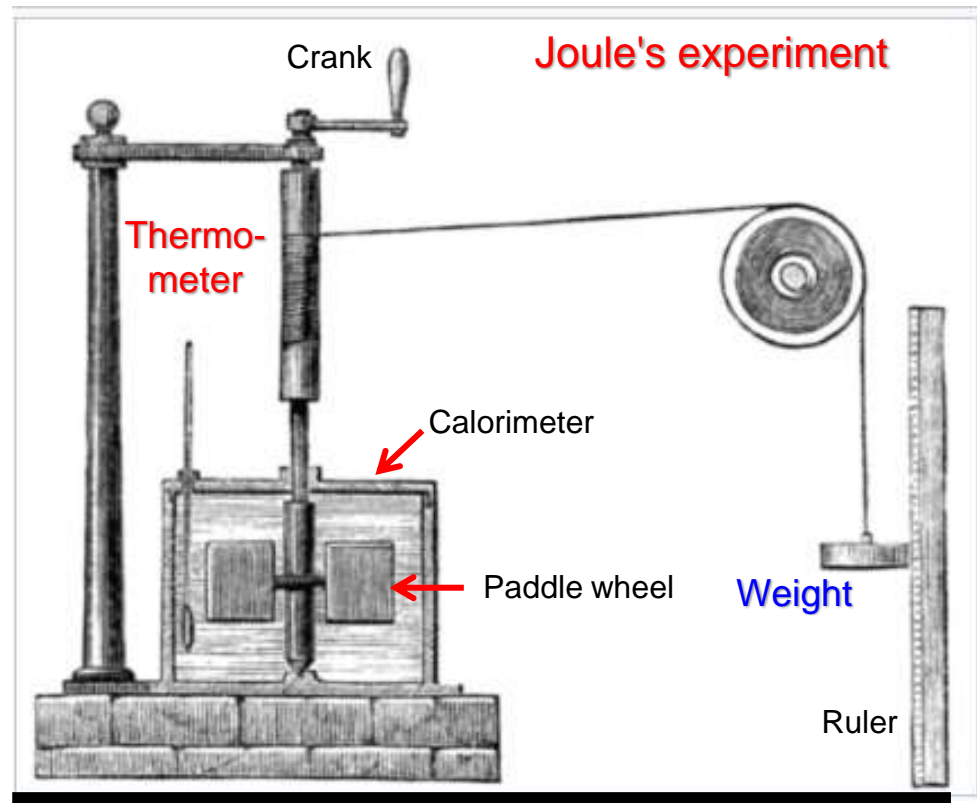
Equivalence of heat and mechanical etc. work ?



Mechanical Equivalent of Heat



"An Experimental Enquiry Concerning the Source of the Heat which is Excited by Friction", (1798), Phil. Transactions of the Royal Society p. 102



Mechanical work equivalent of heat energy

$$W \propto Q \rightarrow W = C \cdot Q$$

$$C = 4.186 \text{ kJ/kcal}$$

Specific heat $C \rightarrow$
Energy (ΔT): $q = C \cdot \Delta T$

Heat energy q required for heating of H_2O : $C_V(\dots) \approx (4.17- 4.22) \text{ J/(g} \cdot ^\circ\text{C)}$
1g material by $\Delta T=1^\circ\text{C}$



Equivalence of Work and Heat Energy

Conservation of total Energy (First Law of Thermodynamics): Energy U of an isolated system is conserved. $\rightarrow U$ is a state function (characterizes state)

If system can exchange energy with environment, it can do that by

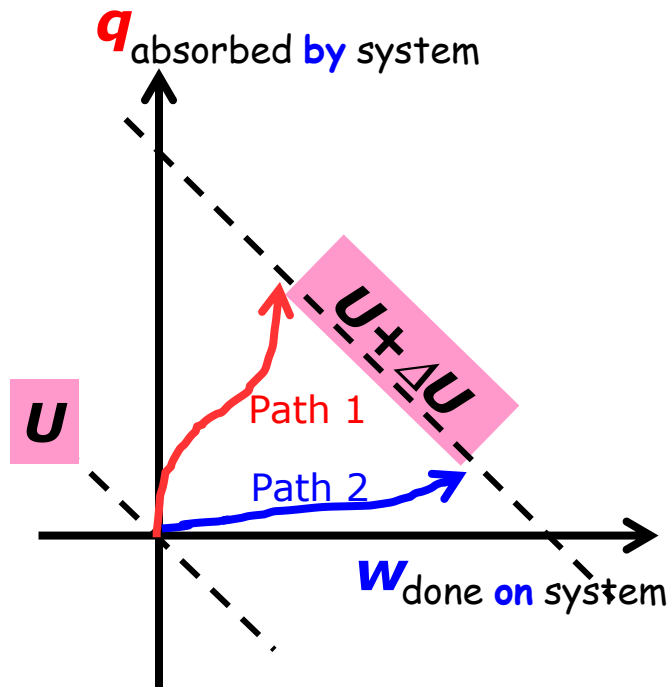
- 1) doing or receiving work w or by
- 2) absorbing or emitting heat q .

$$\Delta U = w + q$$

$\rightarrow W$ and q are path dependent, and NOT state functions.

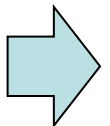
Both are counted positive, if they increase U

Example: Internal energy loss suffered $\Delta U < 0$ = sum of work done by ($w < 0$) and heat emitted by ($q < 0$) by the system.



In asymptotic (long times) state, most likely, system equilibrium state, U **fluctuates in time** and distributes randomly over all degrees of freedom.
 \rightarrow No historic memory about production pathway.

Stationary state of m.p. system, interactions w. environment?



Energy Transfer between *m.p.* Systems & Surroundings

- Energy transferred to and from ***N***-particle systems (e.g., a gas or liquid volume) by mechanical, radiation, or other work is dissipated by multiple interactions of system constituents (particles in a gas or solid lattice).
- Energy dissipation process may take some “relaxation time,” after which all ***N*** system particles move randomly (populate all states) & share total energy on average equally over all degrees of freedom (**Equipartition Law**):

$$\langle \varepsilon_{kin} + \varepsilon_{pot} \rangle_N \propto T$$

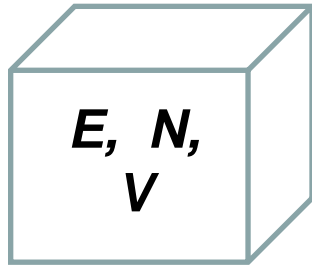
- *Apparently stationary macroscopic state* = “**Thermal Equilibrium**” is a *dynamic process*, with multi-particle configurations fluctuating over entire range with equal probabilities → historic memory is lost (Markov Process).
- **(Total) Energy $U(N)$** is an *extensive state variable* (“state function”), characteristic of the (macroscopic) state of the system, *scales with size*.

$$\langle U \rangle = N \cdot \langle \varepsilon \rangle \propto T \quad (\text{factor} = \text{heat capacity} \quad \Delta \langle U \rangle = C_V \cdot \Delta T)$$

Idealized Classification: Random Statistical Ensembles

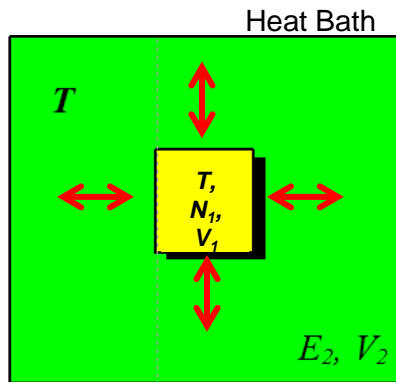
E (or U): total energy, V : containment volume, $N \gg 1$ ($\sim 10^{20}$) constituents

Ensemble: many ($\gg 1$) identically prepared N -particle systems



Micro-canonical ensemble

Isolated systems: Total energy E , linear and angular momentum well defined and conserved. Particle number N and system volume V well defined and conserved. **Macro-& microstates**: E, N, V constant



Canonical ensemble

Closed systems: Particle number N and system volume V well defined and conserved.

Exchange of energy (lin & angular momentum) occurs with surrounding "Heat Bath" \rightarrow fluctuate

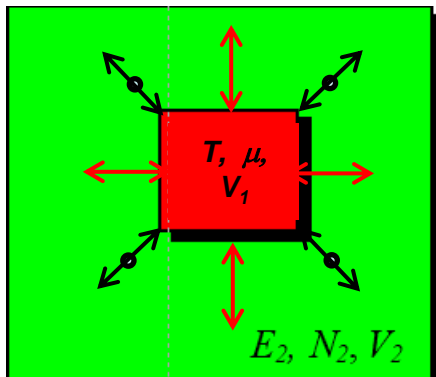
Stationary macro-µ states: T, N, V constant

\rightarrow (System plus reservoirs) = isolated system

Idealized Classification: Random Statistical Ensembles

E (or U): total energy, V : containment volume, $N \gg 1$ ($\sim 10^{20}$) constituents

Ensemble: many ($\gg 1$) identically prepared N -particle systems



Grand-canonical ensemble

Open systems: Exchange of energy and particles occur with surrounding particle reservoirs and "Heat Bath."

Stationary macro- & micro states: T, μ constant.

μ : chemical potential energy per particle.

→ (System plus reservoirs) = isolated system

Heterogenous **system phases** (*solid, liquid, gas*) are homogenous subparts, separating surfaces are **phase boundaries**.

Extensive (additive) state variables scale in proportion to system size (mass, number of particles). Examples: volume, total energy, entropy.

Intensive state variables do not scale with the size of the system.

Examples: pressure, temperature, refractive index

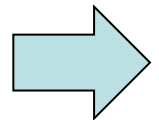
Basic Laws of Statistical Thermodynamics

Statistical Thermodynamics: Physics of equilibration & equilibrium

0) Systems in contact equalize mean internal (kinetic ++) energies,
→ "temperatures" equalize. **Equipartition of thermal energy**
"Canonical & grand canonical" ensembles (of many identical systems).

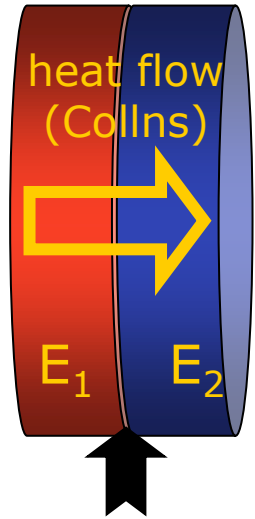
- 1) Energy (of system plus surroundings) is conserved.
- 2) Complex m.p. systems evolve **spontaneously toward state of maximum randomness** (complexity) → Measure = **Entropy** $\Delta S = S_{fin} - S_{in} > 0$
- 3) There exists an absolute "zero point" of energy for any system: energy and complexity are at a minimum (@ **absolute $T=0$ K**).

0. LTD and consequences



Energy Equilibration By Heat Exchange

Composite System
CS = Sys₁ Sys₂



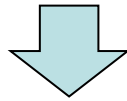
Thermal Contact

Constant energy $E = E_1 + E_2 \rightarrow$ initially in exc. hot $\Omega_1(E_1 = E)$
 \rightarrow Most likely final energy partition E_1/E_2 ?

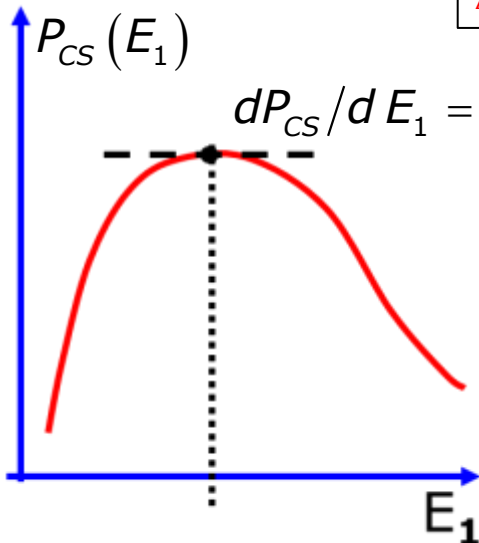
Spontaneous exchange $\Delta q_1 = -\Delta E_1 = +\Delta E_2$ to increase average excited population probability P_{CS} of states in CS

$$\Delta P_{CS}(E, E_1) \propto \Delta \Omega_{CS} = \Delta[\Omega_1(E_1) \cdot \Omega_2(E_2)] \quad \text{maximize } \Omega = \# \text{ populated states}$$

$$= [\Delta \Omega_1(E_1)] \cdot \Omega_2(E_2) + \Omega_1(E_1) \cdot [\Delta \Omega_2(E_2)]$$



$$\frac{\Delta P_{CS}}{\Delta E_1} \propto \frac{\Delta \Omega_1(E_1)}{\Delta E_1} \Omega_2(E_2) - \Omega_1(E_1) \frac{\Delta \Omega_2(E_2)}{\Delta E_2} \quad \text{CS Exc. probability change with } \Delta E_1 = -\Delta E_2 < 0$$



Maximize P_{CS}

$$\frac{dP_{CS}}{dE_1} = 0 \rightarrow$$

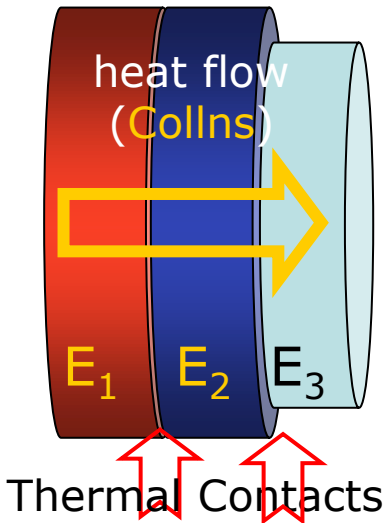
$$\frac{1}{\Omega_1} \frac{d\Omega_1(E_1)}{dE_1} = \frac{1}{\Omega_2} \frac{d\Omega_2(E_2)}{dE_2} \quad \text{dimension } [] = 1/\text{Energy}$$

Depends only on Sys1

Depends only on Sys2

Energy Equilibration By Heat Exchange

Composite System
 $CS = \text{Sys}_1 \text{ Sys}_2 \text{ Sys}_3$



Constant energy $E = E_1 + E_2 + E_3 \rightarrow$ Search for the most likely final energy partition $E_1/E_2/E_3$. For each **spontaneous exchange of heat** $-\Delta q_1 = \Delta E_2 + \Delta E_3 \rightarrow$ increase excited P_{CS}

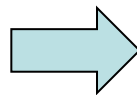
Spontaneous energy transfer *maximizes*

excited $P_{CS} \sim \Omega_{CS} = \Omega_1 \cdot \Omega_2 \cdot \Omega_3$

Check 2nd derivatives (Saddles)

$$\frac{dP_{CS}}{dE_1 dE_2} = 0$$

$$\frac{1}{\Omega_1} \frac{d\Omega_1(E_1)}{dE_1} = \frac{1}{\Omega_2} \frac{d\Omega_2(E_2)}{dE_2} = \frac{1}{\Omega_3} \frac{d\Omega_3(E_3)}{dE_3} = \dots$$

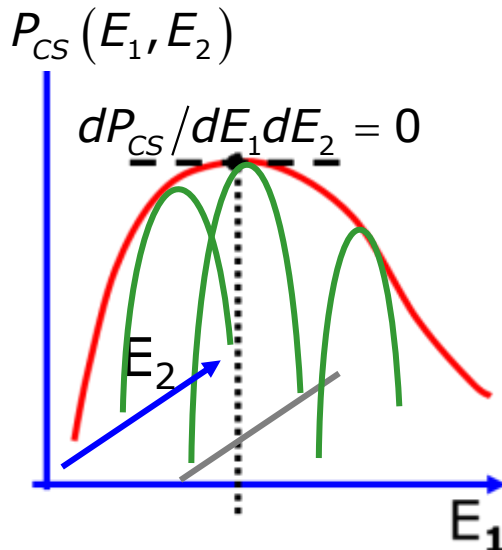


Universal function *dimension* $[] = 1/\text{Energy}$

Independent of index \rightarrow independent of material properties 1,2,...; \rightarrow common mean energy/particle (T)

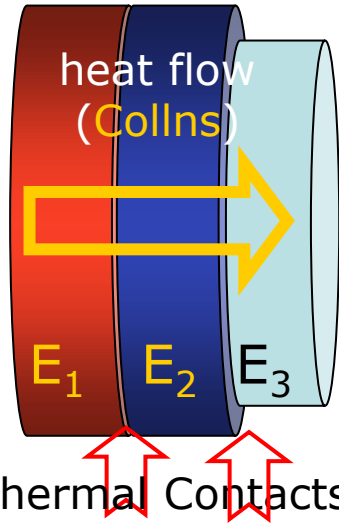
But **common temperature T** depends on total energy E in all systems/total mass \rightarrow

$$\langle E \rangle_N = E/N \sim T$$



Energy Equilibration By Heat Exchange

Composite System
CS=Sys₁ Sys₂ Sys₃



Constant energy $E = E_1 + E_2 + E_3 \rightarrow$ Search for the most likely final energy partition $E_1/E_2/E_3$. For each **spontaneous exchange of heat** $-\Delta q_1 = \Delta E_2 + \Delta E_3 \rightarrow$ increase exc. P_{CS}

Equilibrium partition $\{E_1, \dots, E_n\}$ maximizes excited Ω_{CS}

$$\frac{1}{\Omega_n} \frac{d\Omega_n(E_n)}{dE_n} = \text{const. function } (E = \sum_n E_n)$$

dimension []
= 1/Energy

Plausible Ansatz:

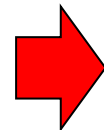
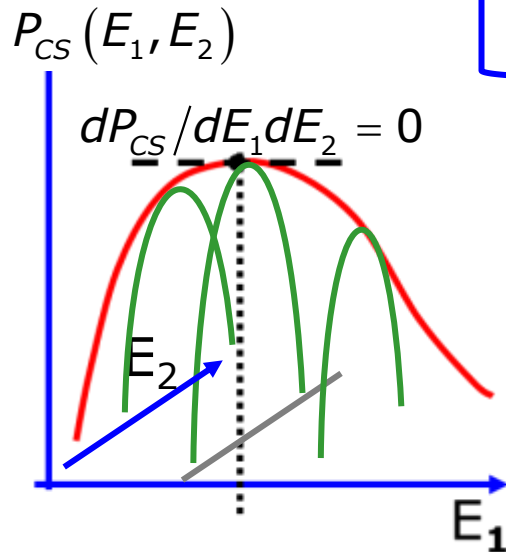
$$\frac{d}{dE} \underbrace{\ln \Omega_{CS}(E)} = \text{const. } (\langle E \rangle) \propto \frac{1}{k \cdot T}$$

$$dS(E)/k \propto \frac{dE}{k \cdot T}$$

Def. = $S/k \rightarrow$ **Entropy** = deg. disorganization

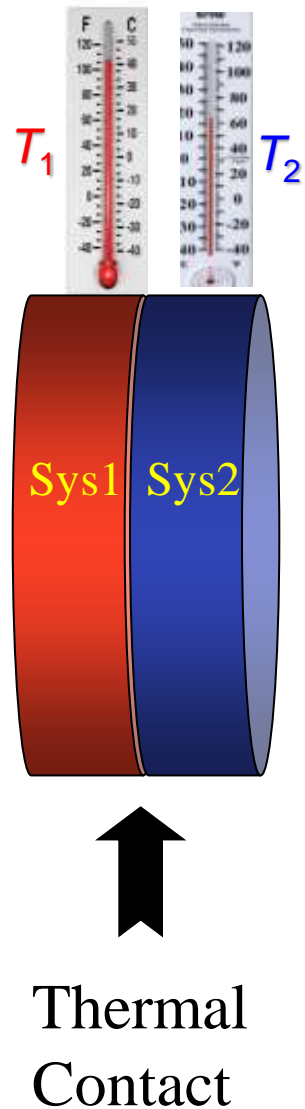
$$dE = T \cdot dS$$

$$dS = \frac{dq}{T} = \frac{\text{Spont. Heat Transfer}}{\text{Temperature}}$$

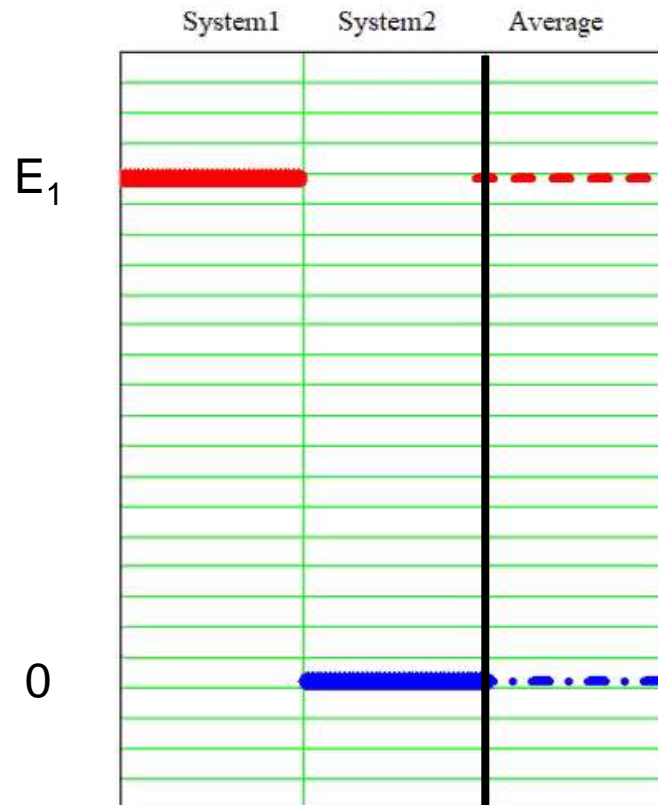


Spontaneous energy redistribution requires entropy increase $\Delta S > 0$ (opposite is not true!)

Simulation: Energy Equilibration



Observation: t -dependent equilibration of temperatures



Discrete energy states of 2 similar **interacting systems**

Initial conditions

$$\langle E_1 \rangle \approx E \quad \langle E_2 \rangle \approx 0$$

Initial mean energies

$$\mathbf{e}_1 = E_1/A_1 > \mathbf{e}_2 = E_2/A_2$$

Random \pm energy and \pm momentum exchange

Final mean energies
thermalization achieved

$$\mathbf{e}_i \rightarrow E_i/A_i \sim T_i = T$$

Equil. energy fluctuations

$$\sigma_{e_i}^2 = k_B T^2 C_V$$

Entropy and Energy at Equilibrium

Independent of material, size, shape, amount → **Intensive function**

Const(E,N)

$$\frac{d}{dE} \text{Ln} \Omega(E, N) = \frac{dS(E, N)}{dE} = \underline{\underline{\text{Const}(E)}}$$

depends on total energy E supplied to one of the sub-systems as heat and the total number of particles. Dimension [Const(E)] = 1/energy.

Obvious *intensive energy* variables: Energy per particle ~ temperature T

→ System energy content of an **ideal gas** (N particles):

$$\langle E \rangle_N = \frac{E}{N} \propto \frac{P \cdot V}{N} = k_B \cdot T$$

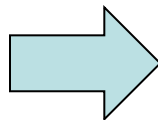
Adopt intensive function S with $\frac{dS}{dE} = \frac{1}{\Omega} \frac{d\Omega(E)}{dE} = \frac{d}{dE} k_B \cdot \text{Ln} \Omega(E) \propto \frac{1}{T}$

Simplest form
compatible
S extensive

Set integration constant = 0

$$S(E) = k_B \text{Ln} \Omega = \cancel{S_0} + \frac{E}{T} \rightarrow \text{Distributed heat energy } E = T \cdot S$$

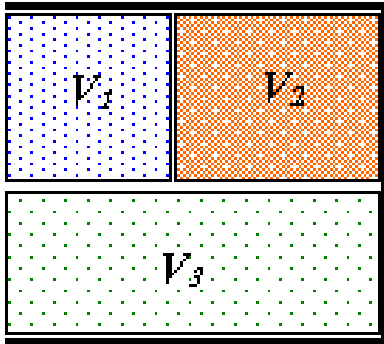
$$\text{Heat absorption / emission by system @T : } \Delta \langle E \rangle = \Delta q \rightarrow \Delta S = \frac{\Delta q}{T}$$



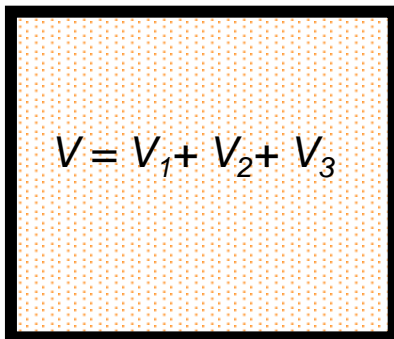
$$\Delta S = S_{fin} - S_{in} = k_B \cdot \text{Ln} \left(\frac{\Omega_{fin}}{\Omega_{in}} \right)$$

Example: Entropy Gain in Gas Mixing

In: Separated Gases



Fin: Mixed Gases



S - gain by individual gas components $\rightarrow \Omega_i \propto V_i$ volume

Entropy is extensive (additive). \rightarrow components add

$$S_i = n_i \cdot R \cdot \ln(V_i) > 0 \quad n_i = \text{number of moles } i$$

$$\text{Total \# moles } n = n = \sum_i n_i; \text{ total vol } V = \sum_i V_i$$

$$\text{Initial } S_{in} = R \cdot \sum_i n_i \cdot \ln(V_i) \rightarrow S_{fin} = R \cdot \sum_i n_i \cdot \ln(V)$$

$$\Delta S = -R \cdot \sum_i n_i \cdot \ln\left(\frac{V_i}{V}\right) > 0 \quad S \nearrow$$

$$\text{total moles } n = \sum_i n_i \rightarrow \Delta S = -n \cdot R \cdot \sum_i \left(\frac{n_i}{n}\right) \cdot \ln\left(\frac{V_i}{V}\right)$$

Relative abundance (probability for) particle type i

$$p_i = \frac{n_i}{n} = \frac{V_i}{V} < 1$$

$$\Delta S = -n \cdot R \cdot \sum_i p_i \cdot \ln p_i \quad n \text{ moles}$$

$$N \text{ particles } \Delta S = -N \cdot k_B \cdot \sum_i p_i \cdot \ln p_i$$

Can use other "State Functions" to calculate entropy, all related by EoS

Entropy of Mixing: Example

One mole each of two **equivalent ideal gases**, **A=O₂** and **B=H₂**, in their respective halves of a separated container at **P_A=P_B= 1atm**, **T_A=T_B= 298K**

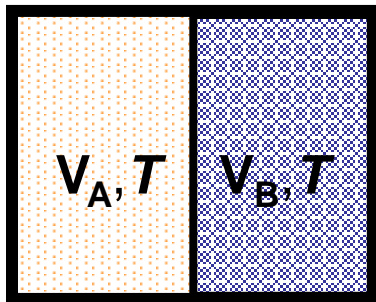
→ s.p. spaces

$$\varpi_A \propto V_A \text{ and } \varpi_B \propto V_B$$

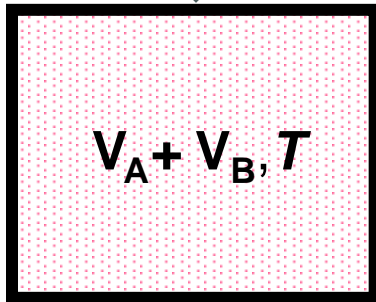
No change in energy

$$\Delta q = 0$$

→ When partition is removed, the gases will mix.



$$\Delta S_{mix} > 0$$



Information on s.p. configuration space from EoS

Ideal gases $P \cdot V = n \cdot R \cdot T \rightarrow \varpi_{A,B} \propto V_{A,B} \propto n_{A,B}$

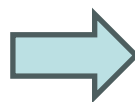
Total **n = 2 moles** ; Mole fractions: $n_A/n = n_B/n = 0.5$.
 → probabilities are $p_A = p_B = 0.5$ also in state space

→ “Entropy of Mixing”

$$\Delta S_{mix} = -nR \left\{ \frac{n_A}{n} \cdot \ln\left(\frac{n_A}{n}\right) + \frac{n_B}{n} \cdot \ln\left(\frac{n_B}{n}\right) \right\} =$$

$$= -2R \{ 0.5 \cdot \ln(0.5) + 0.5 \cdot \ln(0.5) \} = 11.5 \text{ J/mol} \cdot \text{K} > 0$$

Entropy is gained \leftrightarrow mixing occurs spontaneously.



$$\Delta S \geq \frac{\Delta q}{T} \text{ in spontaneous processes}$$

Universal Gas Constant
 $R=8.31 \text{ J/mol} \cdot \text{K}$

Mixed gases are difficult to unmix (costs energy)!

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Maxwell-Boltzmann energy distributions
Fluctuating (Langevin) dissipative forces
- Kinetics of dilute gases
Work and heat transfer
 Flow of heat and radiation
 Laws of thermodynamics, thermodynamic ensembles, entropy
- Fundamental ideal gas laws, Equation of state (EoS)

Reading Assignments

Weeks 5 & 6

LN IV.1-4:

Kondepudi Ch. 1-3.
Additional Material

McQuarrie & Simon
Ch. 2, 5
Math Chapter B, C